Lower-mantle seismic discontinuities and the thermal morphology of subducted slabs

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Abstract

The nature of the lowermost few 100 km of the mantle holds insights into the scale of convection of the Earth’s mantle. Here we investigate the fine structure of this region using seismic migration of shear waves reflecting from velocity discontinuities near the core–mantle boundary (CMB), recorded by a dense network of European seismic stations. Two main features are found. One structure exhibits large-scale topography and can be explained by a sharp increase in seismic velocity 206–316 km above the CMB. The second structure lies 55–85 km above the CMB and marks a sharp reduction in velocity. To explain the origin of these discontinuities, we appeal to numerical simulations of three-dimensional (3D) mantle convection, incorporating rigidly moving plates. The plates organize the convective flow to allow for the development of large convection cells and cool sheet-like downwellings at a convergent plate boundary. The results suggest that, in places, observed seismic complexity in D"_U discontinuity structure may be due to the thermal gradients resulting from the accumulation of cool subducted plate material at the CMB.

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1. Introduction

A long-standing debate in geophysics has revolved around the degree of mass exchange between the upper and lower mantle. Tomographic images show slab structures extending well into the lower mantle, and in some places to the core–mantle boundary (CMB; e.g., [1]), thereby supporting the whole-mantle convection paradigm. However, tomographic results have also formed the basis for arguments for a compositionally distinct layer at the base of the mantle that might explain the source for the primordial geochemical signature observed in oceanic island basalts [2]. Resolving the issue of mass exchange may consequently require an understanding of the properties of the D"_U region, the lowermost 300 km of the mantle.
The anomalous nature of D\\textsuperscript{'} has been recognised for nearly 70 years [3–5]. However, in recent years, seismic studies have illuminated the complexity of this region on a range of length scales through, for example, observations of small-scale structures or scatterers [6], ultralow velocity zones (ULVZs) [7] and seismic anisotropy close to the core–mantle boundary (CMB) [8,9]. It is the site of the largest lateral variations in velocities in the lower mantle [1,10]. There is also abundant evidence for a seismic discontinuity at the top of the D\\textsuperscript{'} region [11,12] and data from seismic arrays have enhanced our images of this discontinuity [13,14]. The complexity of the region has often been used as an argument for chemical heterogeneity at the base of the mantle. Recently, seismic migration, a technique borrowed from oil exploration seismology, has been employed to determine the location of scatterers and the depth of such discontinuities [15–17]. This method tries to find the origin of scattered or reflected waves by back-projecting seismic energy to the likely place of scattering or reflection.

Fig. 1. Sources, receivers and reflection points in the lowermost mantle. The stations include permanent stations of the German Regional Network, stations in the Benelux countries as well as a temporary network across Britain and France (SPICED) [39]. The black stars mark events used in this study, grey stars are events that could not be used due to poor quality signals. The reflection points in the deep mantle (middle figure) are shown for the four events in different grey-scale circles. The ellipses indicate the clusters used in the migration and denote the approximate size of a Fresnel Zone for the seismic waves used in this study. The arrow points to the direction of the station TNS in Germany. The tomographic image is from [1], the grey-scale indicates the S velocity contrast in the lowest layer of the mantle.
2. Data analysis

Here, we have used a migration method [16] to study variations in D\textsuperscript{U} structure beneath northern Asia, looking for reflections from this region. Dense arrays of seismic stations in Europe are used to investigate events beneath the region of the eastern China–Russia border and Japan, yielding source–receiver distance ranges of 68\degree–82\degree. Such data illuminate reflection points in the deep mantle beneath northwestern Russia in an area approximately 10\degree by 20\degree (Fig. 1). This is a region where tomographic S wave models (e.g., [1]) indicate the end of a fast-velocity region, sometimes associated with the ancient Izanagi plate [18]. We look for energy in the coda of S waves that turn just above the D\textsuperscript{U} region. Should there exist a D\textsuperscript{U} discontinuity associated with a sharp increase in velocity, there will be a reflection from the top of the D\textsuperscript{U} region, here simply termed SdS. This wave consists of a reflection from the top of the D\textsuperscript{U} reflector and a diving wave that turns in the D\textsuperscript{U} region. These waves form a triplication in travel time [11]. Given the band-limited nature of seismic data, it is usually difficult to distinguish between the reflected and diving waves and we therefore simply refer to the signal as SdS. It is also possible that such a triplication will be produced by steep velocity gradients in the D\textsuperscript{U} region and not a discontinuity [19]. In order to map detailed variations in the topography of D\textsuperscript{U} discontinuities, we look for this signal using clusters of bounce points roughly the size of a Fresnel zone (220 by 440 km for SdS, with a dominant frequency of 6 s). We have performed the migration for all events and receivers in nine individual and slightly overlapping clusters. The grid spacing is 1\degree (≈ 50 km) laterally and 10\degree vertically and the grid extends 28\degree by 35\degree (1680 by 1500 km) laterally and has 52 depth layers from the core–mantle boundary to 2200 km. Four earthquakes have been studied (Table 1): the number limited by the magnitude of the event, complexities in source functions and low signal to noise ratios.

3. Migration results

An example of the results of migration for data within a single bounce-point cluster is shown in Fig. 2a. The technique starts by calculating travel times between a 3D grid of hypothetical reflection points and individual stations within a cluster. Seismic data from the cluster are then time-shifted by the predicted times and are summed for each hypothetical reflection point. The left-hand side of Fig. 2a shows the results for four different depth slices. The bottom slice (depth of 2888 km) shows coherent positive amplitudes corresponding to S wave energy reflected from the CMB (i.e., the ScS phase). The second layer shows the focusing of negative amplitude energy at a depth of 2825 km. A ring of high amplitudes can be seen in this figure around the centre of the reflection area. These high amplitudes show ScS, the reflection from the CMB, which would arrive several seconds later than SdS at this depth. This is the expected behaviour for migration if the sources and receivers are far apart; the migrated image will focus to a point at the location of a true reflector. If the migration is applied to neighboring depth slices above this point, this image will defocus into circles following the isochrones around this point. The layer at a 2755 km depth shows no focusing of energy, indicating that there is no structure at this depth that produces a reflection. The top slice, 2645 km deep, shows clearly focussed positive amplitude energy.

The results of the migration show a positive amplitude reflection at depths that range between 2545 and 2705 km, depending on the cluster (Fig. 3). This reflection has the same polarity as the lower-mantle turning S wave and the ScS wave and suggests an increase in impedance. We interpret the signal as a reflection from a seismic discontinuity at the top of D\textsuperscript{U}. Other studies have shown evidence for both P and S wave discontinuities in this region and, taken together, they suggest considerable variability in the

<table>
<thead>
<tr>
<th>Event No.</th>
<th>Date</th>
<th>Time</th>
<th>Latitude</th>
<th>Longitude</th>
<th>Depth</th>
</tr>
</thead>
<tbody>
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<td>1</td>
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<td>13:37:02.1</td>
<td>27.89</td>
<td>139.40</td>
<td>498 km</td>
</tr>
<tr>
<td>2</td>
<td>30 March 1995</td>
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<td>44.84</td>
<td>137.54</td>
<td>319 km</td>
</tr>
<tr>
<td>3</td>
<td>31 March 1995</td>
<td>14:01:40.1</td>
<td>38.21</td>
<td>135.01</td>
<td>354 km</td>
</tr>
<tr>
<td>4</td>
<td>8 April 1999</td>
<td>13:10:34.1</td>
<td>43.61</td>
<td>130.35</td>
<td>565 km</td>
</tr>
</tbody>
</table>

Earthquake parameters are from the Preliminary Determination of Earthquakes (PDE) Catalogue.
distance of the reflector from the CMB. An average reflector depth of 2605 [15,20] and 2650 km [17], respectively, for both P and S waves has been observed for this region, but other studies have shown large variations in the depth of the discontinuity [21]. Previous studies have normally used estimates based on inversions for one-dimensional (1D) structure. However, Freybourger et al. [17] also used a migration technique but used a set of European stations to estimate a single reflection point. Our results suggest more dramatic variations in topography (186 to 345 km from the CMB), most likely because we are using a more accurate migration method and clusters of bounce points.

Perhaps a more surprising result is the presence of a weaker reflector 55 to 85 km above the CMB. The
The calculated topography of the second structure is much less than that observed at the top of the $D_U$ region (Fig. 3). Because there is a decrease in impedance, reflections from this boundary are expected to be smaller than those from the upper boundary. Reflections from the positive upper boundary are larger due to near-critical angles of incidence.

**4. Geodynamical constraints**

To interpret our reflections, we appeal to a geodynamical mantle convection model that incorporates thick stiff plates in a 3D Cartesian geometry [22]. We use a hybrid spectral–finite-difference approach to model convection in a $3 \times 2 \times 1$ Cartesian geometry using a grid resolution of $486 \times 324 \times 162$. The velocities are solved using a propagator technique and the energy equation is solved using a flux-corrected finite-difference formulation [23]. We specify periodic (wrap around) boundary conditions on the vertical sidewalls of the calculations. The convection calculation features a depth-dependent viscosity based on a fit of postglacial rebound and free-air gravity anomaly data by Forte and Mitrovica [24], resulting in a viscosity profile that increases smoothly by a factor of 36 from the upper to the lower mantle. Recent work by Forte and Mitrovica [25] suggests that the viscosity of the lower mantle may vary sharply by orders of magnitude at several different depths with several reversals in gradient, but that the ratio of the mean lower mantle viscosity to the mean upper mantle viscosity is still approximately 40. Guided by the persistence in the relevant literature of a mean lower mantle to upper mantle viscosity of between 30 and 40 (e.g., [26–29]), we employ the smooth variation in mantle viscosity employed by Pysklywec and Mitrovica [30] as a first-order approximation of the increase in viscosity with depth through the lower mantle. However, we note that significant rapid increases (and decreases) in mantle viscosity structure could significantly affect our results. The upper 100 km of
our model, the stiff plates, is 1000 times more viscous than the upper mantle immediately below. Similarly, the static plate geometry is specified a priori. However, the time-dependent plate velocities evolve dynamically in accordance with a force-balance method [31]. Consequently, plate motion is determined by the buoyancy distribution within both the plates and the underlying mantle, and the temperature of subducted plate material reflects the history of the plate velocity that is consistent with the other model parameters. Heating is supplied by a combination of an isothermal bottom boundary and uniformly distributed, temporally invariant internal heat sources. The resulting ratio of basal heating to internal heating, as determined by the time-averaged surface heat flow, is approximately 1:1. The Benard–Rayleigh number of the calculation, based on the lower mantle viscosity, is $2.8 \times 10^6$.

Fig. 4a shows two isosurfaces in a temperature field snapshot from a calculation that features two plates. Without the presence of the model plates, an internally heated convecting system in the parameter range we consider is characterized by short wavelength flow in the upper thermal boundary layer of the system (e.g., [32]). The plate size affects the convective wavelength of the system (e.g., [41,42]), and therefore, the temperature of subducted plate material relative to the temperature of fluid below. The convergence of the model plates along the specified linear plate boundary results in the formation of a downwelling sheet of cold material that begins to descend through the mantle, with a morphology and temperature that reflects the linearity of the plate boundary and the size of the plate, respectively. The sheet-like downwelling descends with little deformation until it reaches the lower boundary of the system. In contrast, upwelling material breaks away from the lower thermal boundary layer in sheets but rapidly organizes into cylindrical plume-like features while traversing the deep mantle. Cool material spreads laterally across the base of the system (modelled CMB) upon arrival. There is considerable structure on the upper boundary of this cold isosurface. Moreover, due to variations in plate velocity (and hence subduction rates) sections of cold slab detach from the supply of the downwelling flow, resulting in distinct sheets of cold material accumulating in some regions at the base of the convecting system while other regions may be void of the remnants of cool ancient plates.

An example of the lateral variations in thermal structure of the lowermost mantle is also illustrated in Fig. 4b by showing vertical temperature profiles, for fixed horizontal co-ordinates, through the temperature
field snapshot presented in Fig. 4a. In regions containing the remnants of slab material, there are relatively sharp decreases in temperature a few hundreds of kilometers above the modelled CMB and very sharp increases in temperature much nearer the CMB. Profiles through ambient parts of the temperature field show a simple increase in temperature near the CMB.

We constructed a velocity profile loosely based on thermal profiles for the downwelling regions (Fig. 2b, thick black line) and used the reflectivity method [33] to model wave propagation through the resulting structure.

A rigorous temperature to velocity conversion is unwarranted, given the implicit approximations made in the mantle convection calculations. Some of the most important simplifications specified include the use of a Cartesian geometry, the allowance of two-sided subduction at convergent boundaries, the neglect of a depth-dependent thermal expansion coefficient and the negligence of temperature-dependent viscosity (implying that the slabs in our calculations are strictly thermal features). Moreover, the scaling constants are not well known for the lowermost mantle. Nevertheless, because our calculations incorporate the full thermal history of a dynamically consistent, large, thick plate cooling in a whole-mantle calculation, the thermal signature of the subducted material in our model is more realistic than in previously applied calculations [34].

A sharp increase in velocity gradient at the top of D" will produce a travel time triplication and, hence, ScS precursors. With real data, it can be difficult to distinguish between such a velocity gradient and a discontinuity [19]. A sharp negative gradient will also backscatter low-frequency seismic energy, again producing an apparent seismic reflection. The synthetic waveforms for this model are migrated in the same way as the data, and Fig. 2 shows that there is good agreement between the data and the synthetics for our model.

5. Discussion

The thermal models show large variations in the complexity of the upper boundary of the slab anomaly. The lower boundary, which represents the hot CMB boundary condition, is at a more uniform height. These features are also born out in our observations. The negative-amplitude lower-boundary reflections have not been noticed before, probably due to their weakness—they are not normally visible in the raw seismograms—but Gaherty and Lay [35] reported the observation of a second positive reflection in this region, indicating layering within the D" region. It is interesting to speculate whether or not the lower boundary could generate signals associated with ultralow velocity zones (ULVZs) [7]. However, our reflector is much further from the CMB than the proposed top boundary of ULVZs.

Our results suggest that much of the complexity in discontinuity structure in the lowermost mantle beneath northern Asia can be explained by the thermal structure of a slab that has been subducted to the base of the mantle. This does not preclude the possibility of compositional anomalies in D". It has been suggested that downwelling slab material may bulldoze away primordial material accumulated at the CMB [21]. Comparative variations in shear velocities and bulk-sound-speed agree in slab regions, supporting thermal anomalies but disagreeing low-velocity regions, supporting chemical anomalies [10]. There is also the question of the fate of oceanic crust in the lowermost mantle. Its form and seismic properties at these pressures and temperatures are poorly understood at this time. Inclusions of such material may explain lower-mantle scatterers [16,36] or even seismic anisotropy [8]. However, it can be difficult to explain the seismic properties of D" with a simple chemical layer alone [37].

6. Conclusions

We have migrated ScS precursors to a 3D grid in the lowermost mantle beneath Eurasia detecting two discontinuities. The top reflector at approximately 300 km above the CMB shows an increase in S wave speed and large topography, whereas the second reflector at approximately 70 km above the CMB shows a negative velocity gradient with much less topography. The observed data can be explained using the temperature field of descending thermally defined slabs in simple geodynamic models that simulate the convergence of stiff plates in a 3D mantle convection
calculation, with a heating mode and depth-dependent viscosity in the parameter range appropriate for the Earth. We note that a strong increase in the viscosity stratification or the thermal expansion coefficient of the mantle could affect the descent of the cold material in a convecting system such as that which we modelled. Similarly, we do not model the influence of phase change boundaries on the descending material. Indeed, we consider investigating the influence of each of these complexities on the depth to which plates descend beyond the scope of this work. Rather, we aim to determine whether a thermal anomaly with a plate-like signature could explain our seismic observations in the case that the plate material does reach the CMB.

Our results are another piece of evidence that slabs may reach the CMB and that convection can operate on a whole-mantle scale. Our geodynamical simulations show that as slab-like thermal anomalies reach the CMB, they can emplace thermal features that can explain the origin of the seismic properties of the lowermost mantle in the regions sampled by our observations.

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References

[18] C. Lithgow-Bertelloni, M.A. Richards, The dynamics of Ce-


